



# **CORIOLIS COUPLING IN A HÉNON-HEILES SYSTEM**

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## THE SYSTEM

**Two-degrees of freedom rotating system** with Hamiltonian

$$\mathcal{H} = \frac{1}{2}(X^2 + Y^2) \frac{-\omega(xY - yX)}{-\omega(xY - yX)} + \frac{1}{2}(x^2 + y^2) + yx^2 - \frac{1}{3}y^3$$
  
Coriolis term

- (x, y) cartesian coordinates and (X, Y) conjugate canonical momenta
- We use a **co-rotating reference** frame
- $\omega$  parameter: the rotational angular velocity of the system
- We study the effect of the angular velocity  $\omega$  in the escape dynamics

Hamiltonians with Coriolis term appear in contexts such as:

- Atomic and molecular Physics
- Celestial mechanics
- Galactic dynamics

### **EQUILIBRIUM POINTS**

**Four equilibrium points:**  $(x_e, y_e, X_e, Y_e)$  $E_0 \equiv (0, 0, 0, 0)$  $E_1 \equiv (0, 1 - \omega^2, \omega(\omega^2 - 1), 0)$  $E_2 \equiv \left(\frac{\sqrt{3}}{2}(1-\omega^2), \frac{1}{2}(\omega^2-1), \frac{1}{2}\omega(1-\omega^2), \frac{\sqrt{3}}{2}\omega(1-\omega^2)\right)$  $E_3 \equiv \left(\frac{\sqrt{3}}{2}(\omega^2 - 1), \frac{1}{2}(\omega^2 - 1), \frac{1}{2}\omega(1 - \omega^2), \frac{\sqrt{3}}{2}\omega(\omega^2 - 1)\right)$ located in the x = Y = 0 manifold  $E_1$ located symmetrical respect to that manifold  $E_{2.3}$ 

Stability and energy of the equilibria:

 $E_0 \qquad \text{Stable equilibrium with energy} \quad \varepsilon_0 = 0$  $E_{1,2,3} \qquad \text{Unstable center} \times \text{saddle equilibrium points}$ with the same energy  $\varepsilon_{1,2,3} = \frac{(1-\omega)^3}{6}$ 

EFFECTIVE POTENTIAL U(x, y). EVOLUTION WITH  $\omega$ 

$$U(x,y) = \mathcal{H} - \frac{1}{2}(\dot{x}^2 + \dot{y}^2) = \frac{1 - \omega^2}{2}(x^2 + y^2) + x^2y - \frac{1}{3}y^3$$



#### **INNER TRAPPING REGION AND ESCAPE CHANNELS**



The orbits are **confined** in a inner trapping region around the origin

The orbits can **escape** from the inner región through **3 different channels.** 

We study the escape dynamics for an energy  $\mathcal{E} = 1.2 \mathcal{E}_{saddles}$ 

#### SURFACES OF SECTION. EVOLUTION WITH $\omega$



#### LONG LIVING STICKY ORBITS







## ESCAPE PROBABILITY. EVOLUTION WITH $\boldsymbol{\omega}$



Escape probability computed from the escape basins Escape probability computed from all the phase space volume

## CONCLUSIONS

Effects of the rotational angular velocity  $\omega$  on the escape dynamics as the angular velocity  $\omega$  increases from 0 to 1:

- The size of the inner trapping region shrinks.
- The chaoticity and uncertainty of the escape dynamics drecreases.
- The escape probability also drecreases non-monotonically.